

MIT OCW GR PSET 1

- 1 a. Show that the sum of any two orthogonal spacelike vectors is spacelike

"spacelike" $\Rightarrow \vec{A} \cdot \vec{A} > 0, \vec{B} \cdot \vec{B} > 0$

"orthogonal" $\Rightarrow \vec{A} \cdot \vec{B} = 0$

$$\vec{A} + \vec{B} \equiv \vec{C}$$

$$\Rightarrow \vec{C} \cdot \vec{C} = \underbrace{\vec{A} \cdot \vec{A}}_{>0} + \underbrace{\vec{B} \cdot \vec{B}}_{>0} + \underbrace{2\vec{A} \cdot \vec{B}}_0$$

$\Rightarrow \vec{C} \cdot \vec{C} > 0$ and $\vec{A} + \vec{B}$ is also spacelike

- b. Show that a timelike vector and a null vector cannot be orthogonal

If \vec{A} timelike, then $\vec{A} \cdot \vec{A} < 0$ $\vec{B} \neq 0$

If \vec{B} null, then $\vec{B} \cdot \vec{B} = 0$ and ~~timelike~~

Assume $\vec{A} \cdot \vec{B}$ is orthogonal i.e. $\vec{A} \cdot \vec{B} = 0$

$$-(B^0)^2 + |\vec{B}|^2 = 0$$

$$\Rightarrow (B^0)^2 = |\vec{B}|^2$$

$$-A^0B^0 + A^1B^1 + A^2B^2 + A^3B^3 = 0$$

• choose a frame where \vec{A} is completely temporal : $\vec{A} = (A^0, 0, 0, 0)$

$$\Rightarrow A^0B^0 = 0, \text{ Also } (A^0)^2 < 0$$

$$\Rightarrow B^0 = 0$$

$$\Rightarrow (B^0)^2 = |\vec{B}|^2 = 0$$

$$\Rightarrow (B^1)^2 + (B^2)^2 + (B^3)^2 = 0$$

$$\Rightarrow B^1 = B^2 = B^3 = 0$$

$$\Rightarrow \vec{B} = (0, \vec{0})$$

• which contradicts our definition of a null or "light like" vector \vec{B} and implies

$$\vec{A} \cdot \vec{B} \text{ cannot } = 0 \quad \times$$

2. In some reference frame \vec{J} and \vec{D} have the components :

$$J^{\pm} \doteq (1+t^2, t^2, \sqrt{2}t, 0)$$

$$D^{\pm} \doteq (x, stx, \sqrt{2}t, 0)$$

$\vec{A} \cdot \vec{B}$ is Lorentz invariant
so we are free to work in this frame w.l.o.g.

and the scalar $\rho = x^2 + t^2 - y^2$

a. Show that \vec{U} is suitable as a 4-velocity.
Is \vec{D} also?

a 4-velocity requires that $\vec{U} \cdot \vec{U} = -1$

$$\begin{aligned}\vec{U} \cdot \vec{U} &= -(1+t^2)^2 + t^4 + 2t^2 \\ &= -1 - t^4 - 2t^2 + t^4 + 2t^2 \\ &= \boxed{-1} \quad \checkmark\end{aligned}$$

What about \vec{D} ?

$$\begin{aligned}\vec{D} \cdot \vec{D} &= -x^2 + 5^2 t^2 x^2 + 2t^2 \\ &\neq -1 \quad \boxed{\text{so } \text{NO}}\end{aligned}$$

b. Find the spatial velocity \vec{V} of a particle whose 4-velocity is \vec{U} , for arbitrary t :

$$\vec{U} \cdot \vec{U} = -(v^0)^2 + |\vec{V}|^2$$

$$\Rightarrow |\vec{V}|^2 = (v^0)^2 - 1 = 1 + 2t^2 + t^4 - 1$$

$$\Rightarrow |\vec{V}| = \sqrt{2t^2 + t^4} = \boxed{t \sqrt{2+t^2}} = |\vec{V}|$$

$$\left. \lim_{t \rightarrow 0} |\vec{V}| = 0, \lim_{t \rightarrow \infty} |\vec{V}| \rightarrow \infty \approx t^2 \right]$$

C · Find $\partial_\beta U^2$ for all α, β . Show that $U_2 \partial_\beta U^2 = 0$ via brute force

α	β	$\partial_\beta U^2$
t	t	$2t$
t	x	0
t	y	0
t	z	0
x	t	$2t$
x	x	0
x	y	0
x	z	0
y	t	$\sqrt{2}t$
y	x	0
y	y	0
y	z	0
z	t	0
z	x	0
z	y	0
z	z	0

· one could cleverly show that:

$$\begin{aligned} U_2 \partial_\beta U^2 &= \partial_\beta U_2 U^2 \\ &= \partial_\beta (\vec{U} \cdot \vec{U}) \\ &= \partial_\beta (-1) = 0 \quad \checkmark \end{aligned}$$

· But we will use brute force:

$$\begin{aligned} U_2 \partial_\beta U^2 &= -U_t \partial_t U^t \\ &\quad + U_x \partial_t U^x \\ &\quad + U_y \partial_t U^y \\ &\quad + U_z \partial_t U^z \\ &= -(1+t^2)(2t) \\ &\quad + (t^2)(2t) \\ &\quad + (\sqrt{2}t)(\sqrt{2}) \\ &= 0 \end{aligned}$$

$\Rightarrow U_2 \partial_\beta U^2 = 0 \quad \checkmark$

d Find $\partial_{\alpha} D^{\alpha} \rightarrow$ represents a set of 4 numbers that we sum together?

$$\begin{aligned}\partial_{\alpha} D^{\alpha} &= +\partial_t D^t + \partial_x D^x + \partial_y D^y + \partial_z D^z \\ &= 0 + St + 0 + 0\end{aligned}$$

$$\Rightarrow \boxed{\partial_{\alpha} D^{\alpha} = St}$$

e Find $\partial_{\beta} (U^{\alpha} D^{\beta})$ for all α

$$= U^{\alpha} \partial_{\beta} D^{\beta} = \boxed{St U^{\alpha}} \quad -\text{or equivalently}-$$

$$= (St + St^3, St^4, St^2, 0)$$

f Find $U_{\alpha} \partial_{\beta} (U^{\alpha} D^{\beta})$

$$= St U_{\alpha} U^{\alpha} = St (\vec{U} \cdot \vec{U}) = \boxed{-St}$$

This is similar to d since the expressions are actually equivalent up to a minus sign $\cancel{\wedge}$

g calculate $\partial_{\alpha} P$ for all α . Calculate $\partial^{\alpha} P$:

$$\partial_2 \mathcal{P} \rightarrow \{\partial_t \mathcal{P}, \partial_x \mathcal{P}, \partial_y \mathcal{P}, \partial_z \mathcal{P}\}$$

$$\partial_2 \mathcal{P} = \{zt, zx, -zy, 0\}$$

$\partial^2 \mathcal{P}$ is shorthand for:

$$\partial^2 \mathcal{P} = g^{\alpha\beta} \partial_\beta \mathcal{P} \quad \text{where } g^{\alpha\beta} \text{ is the metric tensor} = \text{diag } (-1, 1, 1, 1)$$

therefore,

$$\partial^2 \mathcal{P} = \begin{bmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} zt \\ zx \\ -zy \\ 0 \end{bmatrix} = \begin{bmatrix} -zt \\ zx \\ -zy \\ 0 \end{bmatrix} = \partial^2 \mathcal{P}$$

h) Find $\nabla_{\vec{U}} \mathcal{P}$ and $\nabla_{\vec{D}} \mathcal{P}$:

$$\nabla^2 \partial_2 \mathcal{P} \quad \nabla^2 \partial_2 \mathcal{P}$$

$$zt + 2t^4 + 2xt^2 - 2\sqrt{2}yt$$

$$ztx + 10x^2t - 2\sqrt{2}yt$$

group like terms

3. Consider a timelike unit 4-vector \vec{U} and the tensor:

$$P_{\alpha\beta} = n_{\alpha\beta} + U_\alpha U_\beta$$

a. Show that $V_\perp^\alpha = P_{\beta}^{\alpha} V^\beta$ is orthogonal

to \vec{U} :

We know $\vec{U} \cdot \vec{U} < 0$ "timelike" and unit vector

$$\text{so } \Rightarrow \vec{U} \cdot \vec{U} = -1$$

Let $P_{\beta}^{\alpha} = n_{\beta}^{\alpha} + U^{\alpha} U_{\beta}$

$$\Rightarrow P_{\beta}^{\alpha} V^\beta = n_{\beta}^{\alpha} V^\beta + U^{\alpha} U_{\beta} V^\beta \equiv V_\perp^\alpha$$

$$\Rightarrow U_\alpha V_\perp^\alpha = \underbrace{U_\alpha n_{\beta}^{\alpha} V^\beta}_{\vec{U} \cdot \vec{V}} + \underbrace{U_\alpha U^{\alpha} U_{\beta} V^\beta}_{-1 \quad \vec{U} \cdot \vec{V}}$$

$$\Rightarrow U_\alpha V_\perp^\alpha = \vec{U} \cdot \vec{V}_\perp = \vec{U} \cdot \vec{V} - \vec{U} \cdot \vec{V} = 0$$

$$\Rightarrow \boxed{\vec{U} \cdot \vec{V}_\perp = 0 \quad \checkmark}$$

b) Show that $V_{\perp}^{\perp} = P_{\beta}^{\perp} V^{\beta}$ is unaffected by further projections:

i.e. Show that $V_{\perp\perp}^{\perp} = P_{\beta}^{\perp} V_{\perp}^{\beta} = V_{\perp}^{\perp}$

$$P_{\beta}^{\perp} V_{\perp}^{\beta} = h_{\beta}^{\perp} V_{\perp}^{\beta} + \underbrace{U^{\perp} U_{\beta} V_{\perp}^{\beta}}_{\vec{U} \cdot \vec{V}_{\perp} = \emptyset} = V_{\perp\perp}^{\perp}$$

$$= h_{\beta}^{\perp} V_{\perp}^{\beta} = V_{\perp}^{\perp}$$

$$\Rightarrow \boxed{V_{\perp\perp}^{\perp} = V_{\perp}^{\perp} \quad \checkmark}$$

c) Show that $P_{\alpha\beta} V_{\perp}^{\perp} W_{\perp}^{\beta} = \vec{V}_{\perp} \cdot \vec{W}_{\perp}$

$$\text{RHS} = \vec{V}_{\perp} \cdot \vec{W}_{\perp} = V_{\perp}^{\perp} W_{\perp\perp}$$

$$= P_{\beta}^{\perp} V^{\beta} P_{\alpha}^{\gamma} W_{\gamma} = V^{\beta} W_{\gamma} (n_{\beta}^{\perp} + U^{\perp} U_{\beta}) (n_{\alpha}^{\gamma} + U^{\gamma} U_{\alpha})$$

$$= V^{\beta} W_{\gamma} \delta_{\beta}^{\gamma} + V^{\beta} W_{\gamma} U^{\perp} U_{\alpha} V_{\alpha}^{\gamma} U^{\gamma}$$

$$+ V^{\beta} W_{\gamma} U^{\perp} U_{\beta} n_{\alpha}^{\gamma} + V^{\beta} W_{\gamma} W_{\beta}^{\perp} V^{\gamma} U_{\alpha}$$

$$= V^{\beta} W_{\beta} - (\vec{V} \cdot \vec{V})(\vec{V} \cdot \vec{W}) + 2(\vec{V} \cdot \vec{V})(\vec{V} \cdot \vec{W})$$

$$= (\vec{V} \cdot \vec{W}) + (\vec{V} \cdot \vec{V})(\vec{V} \cdot \vec{W}) \quad \checkmark \quad \sim \sim \sim$$

$$\begin{aligned}
 LHS &= P_{\alpha\beta} V_\perp^\alpha W_\perp^\beta \\
 &= (n_{\alpha\beta} + v_\alpha v_\beta) (P_\gamma^\alpha V^\gamma) (P_\chi^\beta W^\chi) \\
 &= (n_{\alpha\beta} + v_\alpha v_\beta) (n_\gamma^\alpha + v^\alpha v_\gamma) (n_\chi^\beta + v^\beta v_\chi) V^\gamma V^\chi \\
 &= (n_{\alpha\beta} n_\gamma^\alpha + v_\alpha v_\beta n_\gamma^\alpha + n_{\alpha\beta} v^\alpha v_\gamma + v_\alpha v_\beta v^\alpha v_\gamma) \\
 &\quad \cdot (n_\chi^\beta + v^\beta v_\chi) V^\gamma W^\chi
 \end{aligned}$$

$$\begin{aligned}
 &= n_{\alpha\beta} n_\gamma^\alpha n_\chi^\beta + v_\alpha v_\beta n_\gamma^\alpha n_\chi^\beta V^\gamma W^\chi \\
 &\quad + n_{\alpha\beta} n_\chi^\beta v^\alpha v_\gamma V^\gamma W^\chi + v_\alpha v_\beta v^\alpha v_\gamma n_\chi^\beta V^\gamma W^\chi \\
 &\quad + n_{\alpha\beta} n_\gamma^\alpha v^\beta v_\chi V^\gamma W^\chi + v_\alpha v_\beta v^\beta v_\chi n_\gamma^\alpha V^\gamma W^\chi \\
 &\quad + n_{\alpha\beta} v^\alpha v_\gamma v^\beta v_\chi V^\gamma W^\chi + v_\alpha v_\beta v^\alpha v_\gamma v^\beta v_\chi V^\gamma W^\chi \\
 &= (\vec{v} \cdot \vec{w}) + (\vec{v} \cdot \vec{v})(\vec{v} \cdot \vec{w}) + \cancel{(\vec{v} \cdot \vec{v})(\vec{v} \cdot \vec{w}) - (\vec{v} \cdot \vec{v}) \times (\vec{v} \cdot \vec{w})} \\
 &\quad \cancel{+ (\vec{v} \cdot \vec{v})(\vec{v} \cdot \vec{w}) - (\vec{v} \cdot \vec{v}) \cancel{(\vec{v} \cdot \vec{w})}} - \cancel{(\vec{v} \cdot \vec{v})} \cancel{(\vec{v} \cdot \vec{w})} \\
 &\quad + (\vec{v} \cdot \vec{v})(\vec{v} \cdot \vec{w}) = (\vec{v} \cdot \vec{w}) + (\vec{v} \cdot \vec{v})(\vec{v} \cdot \vec{w}) \quad \text{W}
 \end{aligned}$$

$$\Rightarrow \text{LHS} = \text{RHS} \Rightarrow P_{\alpha\beta} V_1^\alpha W_\perp^\beta = \vec{V}_1 \cdot \vec{W}$$

[d] Show that for an arbitrary non-null vector \vec{q} , the projection tensor is given by:

$$P_{\alpha\beta}(q^\alpha) = n_{\alpha\beta} - \frac{q_\alpha q_\beta}{q^r q_r}$$

If this is the case, then $q_\perp = n_{\alpha\beta} - \frac{q_\alpha q_\beta}{q^r q_r}$
and q^\perp should be orthogonal:

$$q_\perp^\alpha q^\perp = n_{\alpha\beta} q^\perp - \frac{q_\alpha q^\perp q^\beta}{q^r q_r}$$

$$= q_\beta - q_\beta = 0$$

It doesn't make sense to have a projection tensor for null vectors since every vector is orthogonal to a null vector ...

- 4 Let $\Lambda_B(\bar{v})$ be a Lorentz boost associated with
3 velocity \bar{v} . Consider

$$\Lambda_{\text{tot}} = \Lambda_B(\bar{v}_1) \Lambda_B(\bar{v}_2) \Lambda_B(-\bar{v}_1) \Lambda_B(-\bar{v}_2)$$

where $\bar{v}_1 \cdot \bar{v}_2 = 0$ and $v_1, v_2 \ll 1$. Show that Λ_{tot} is
a rotation. What is the axis + angle.

- Intuitively it makes sense that the composition of
these 4 boosts should generally get you back to
the same frame, but since they don't commute,
there is some mixing going on that causes these
rotations.
- Take the case:

$$\Lambda_{\text{tot}} = \Lambda_B(v_x) \Lambda_B(v_y) \Lambda_B(-v_x) \Lambda_B(-v_y)$$

where $v_x \perp v_y$

$$\Lambda_B(v_x) = \begin{bmatrix} \gamma_1 & -\gamma_1 \beta_1 & 0 & 0 \\ -\gamma_1 \beta_1 & \gamma_1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

and 

$$\Lambda_B(v_y) = \begin{bmatrix} \gamma_2 & 0 & -\gamma_2 \beta_2 & 0 \\ 0 & 1 & 0 & 0 \\ -\gamma_2 \beta_2 & 0 & \gamma_2 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

and for $\Lambda_B(-v_x)$ and $\Lambda_B(-v_y)$ they are similar
but all the entries are positive

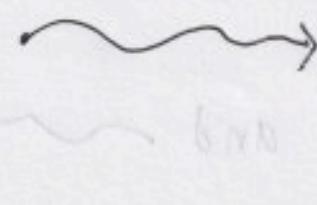
$$\Rightarrow \Lambda_B(-v_x) \Lambda_B(-v_y) = \begin{bmatrix} +\gamma_1 & +\gamma_1 \beta_1 & 0 & 0 \\ +\gamma_1 \beta_1 & +\gamma_1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \gamma_2 & 0 & \gamma_2 \beta_2 & 0 \\ 0 & 1 & 0 & 0 \\ \gamma_2 \beta_2 & 0 & \gamma_2 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

$$= \begin{bmatrix} \gamma_1 \gamma_2 & \gamma_1 \beta_1 & \gamma_1 \gamma_2 \beta_2 & 0 \\ \gamma_1 \gamma_2 \beta_1 & \gamma_1 & \gamma_1 \gamma_2 \beta_1 \beta_2 & 0 \\ \gamma_2 \beta_2 & 0 & \gamma_2 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

Also

$$\Lambda_B(v_x) \Lambda_B(v_y) = \begin{bmatrix} \gamma_1 & -\gamma_1 \beta_1 & 0 & 0 \\ -\gamma_1 \beta_1 & \gamma_1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \gamma_2 & 0 & -\gamma_2 \beta_2 & 0 \\ 0 & 1 & 0 & 0 \\ -\gamma_2 \beta_2 & 0 & \gamma_2 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

$$= \begin{bmatrix} \gamma_1 \gamma_2 & -\gamma_1 \beta_1 & -\gamma_1 \gamma_2 \beta_2 & 0 \\ -\gamma_1 \gamma_2 \beta_1 & \gamma_1 & \gamma_1 \gamma_2 \beta_1 \beta_2 & 0 \\ -\gamma_2 \beta_2 & 0 & \gamma_2 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$



$$\Lambda_{tot} = \Lambda(v_x) \Lambda(v_y) \Lambda(-v_x) \Lambda(-v_y)$$

$$= \begin{bmatrix} \gamma_1 \gamma_2 & -\gamma_1 \beta_1 & -\gamma_1 \gamma_2 \beta_2 & 0 \\ -\gamma_1 \gamma_2 \beta_1 & \gamma_1 & \gamma_1 \gamma_2 \beta_1 \beta_2 & 0 \\ -\gamma_2 \beta_2 & 0 & \gamma_2 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \gamma_1 \gamma_2 & \gamma_1 \beta_1 & \gamma_1 \gamma_2 \beta_2 & 0 \\ \gamma_1 \gamma_2 \beta_1 & \gamma_1 & \gamma_1 \gamma_2 \beta_1 \beta_2 & 0 \\ \gamma_2 \beta_2 & 0 & \gamma_2 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

$$= \boxed{\begin{bmatrix} \gamma_1^2 \gamma_2^2 - \gamma_1^2 \gamma_2^2 \beta_1^2 - \gamma_1^2 \gamma_2^2 \beta_2^2 & -\gamma_1^2 \beta_1 & \gamma_1^2 \gamma_2^2 \beta_2 - \gamma_1^2 \gamma_2^2 \beta_1 \beta_2 - \gamma_1^2 \gamma_2^2 \beta_2^2 & 0 \\ -\gamma_1^2 \gamma_2^2 \beta_1 + \gamma_1^2 \gamma_2^2 \beta_1 + \gamma_1^2 \gamma_2^2 \beta_1 \beta_2 & -\gamma_1^2 \gamma_2^2 \beta_1^2 + \gamma_1^2 & -\gamma_1^2 \gamma_2^2 \beta_2 \beta_1 + \gamma_1^2 \gamma_2^2 \beta_1 \beta_2 + \gamma_1^2 \gamma_2^2 \beta_1 \beta_2 & 0 \\ -\gamma_1^2 \gamma_2^2 \beta_2 + \gamma_2^2 \beta_2 & -\gamma_1 \gamma_2 \beta_1 \beta_2 & \gamma_2^2 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}}$$

This huge annoying 4×4 matrix should in theory represent a rotation around z . Therefore, we compare the circled 3×3 matrix to a rotation matrix.

$$\begin{bmatrix} -\gamma_1^2 \gamma_2^2 \beta_1^2 + \gamma_1^2 & -\gamma_1^2 \gamma_2^2 \beta_1 \beta_2 + \gamma_1^2 \gamma_2^2 \beta_1 \beta_2 + \gamma_1^2 \gamma_2^2 \beta_1 \beta_2 & 0 \\ -\gamma_1 \gamma_2 \beta_1 \beta_2 & \gamma_2^2 & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

$$\stackrel{?}{=} \begin{bmatrix} \cos \theta & -\sin \theta & 0 \\ \sin \theta & \cos \theta & 0 \\ 0 & 0 & 1 \end{bmatrix} \quad \rightsquigarrow$$

• That would imply:

$$\sin \theta = -\gamma_1 \gamma_2 \beta_1 \beta_2$$

$\approx \frac{-v_1 v_2}{\sqrt{(1-v_1^2)(1-v_2^2)}}$

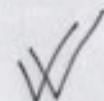
units with $c=1$

denominator ≈ 1 since
 $v_1 \ll 1$ and $v_2 \ll 1$

$\Rightarrow |\theta| \approx v_1 v_2$ about the z axis.

Back in units with $c=c$, we get

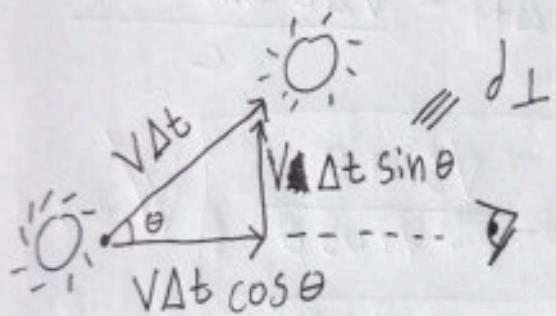
$$|\theta| \approx \frac{v_1 v_2}{c^2} \quad \text{about z-axis}$$



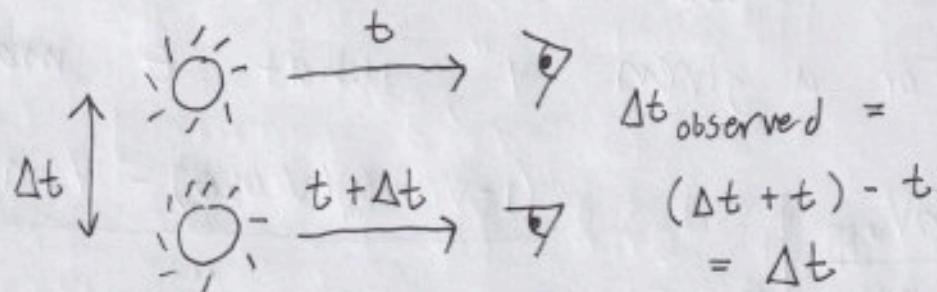
5. A quasar is moving towards you and up at an angle θ . The apparent upwards velocity is v_{app} :

a) Find the expression for v_{app} in terms of θ and the true velocity "v":

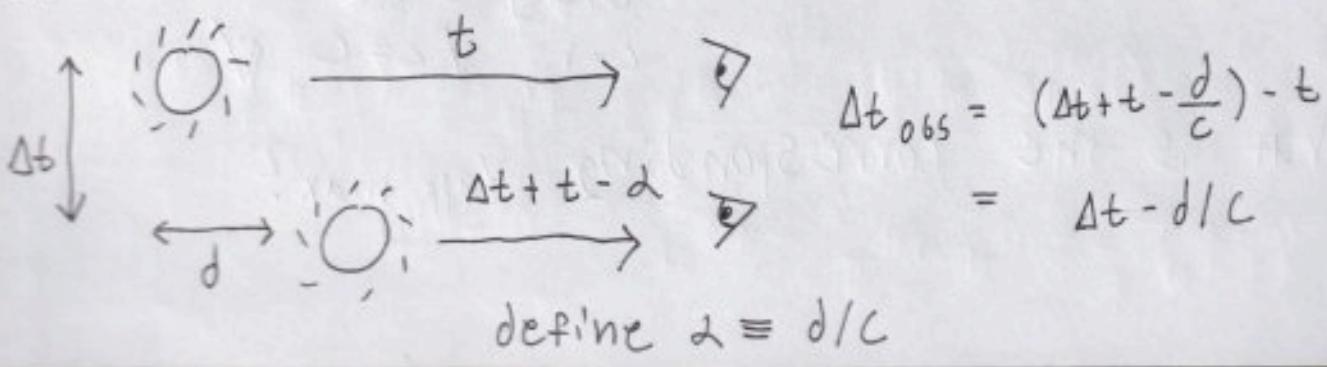
• Let's draw a diagram:



• Also take an aside & start to think about light discretely. Imagine a light bulb releasing photons at intervals of Δt :



• If the light bulb is stationary, then you also perceive the time between flashes as Δt as above. However, if the bulb is moving towards you:



- Then $\Delta t_{\text{obs}} = \Delta t - d/c \neq \Delta t$
i.e. you see a different time between emission of photons compared to the proper time of the bulb.
- We can say that the perceived upwards velocity of the quasar is:

$$V_{\text{app}} = \frac{d_L}{\Delta t_{\text{obs}}} ; \quad \Delta t_{\text{obs}} = \Delta t - \frac{d_{\parallel}/c}{c} \\ = \Delta t - \frac{\sqrt{\Delta t^2 - c^2} \cos \theta}{c}$$

$$\Rightarrow V_{\text{app}} = \frac{\sqrt{\Delta t^2 - c^2} \sin \theta}{\Delta t - \sqrt{\Delta t^2 - c^2} \cos \theta} \Rightarrow V_{\text{app}} = \frac{\sqrt{v^2 - 1} \sin \theta}{1 - \sqrt{v^2 - 1} \cos \theta} \checkmark$$

- b. For a given "v", what θ maximizes V_{app} ?

$$0 = \frac{\partial V_{\text{app}}}{\partial \theta} \Big|_{\theta_{\text{max}}} = \frac{(1 - v \cos \theta_m)(v \cos \theta_m) - v \sin \theta_m(v \sin \theta_m)}{\dots}$$

$$\Rightarrow v \cos \theta_{\text{max}} - v^2(\sin^2 \theta_{\text{max}} + \cos^2 \theta_{\text{max}}) = 0$$

$$\Rightarrow \cos \theta_{\text{max}} = v \Rightarrow \theta_{\text{max}} = \cos^{-1}(v)$$

Okay since in units with $c=1$, $\sqrt{c}=c$ \checkmark

- What is the corresponding $V_{\text{app}, \text{max}}$?

$$V_{am} = \frac{\sqrt{v} \sin(\cos^{-1}(v))}{1-v^2}, \quad v < 1$$

↑ bounded to $v=1$
 ↓ really small if $v \ll 1$

→ possible for $V_{am} > 1$ (in units with $c=1$)

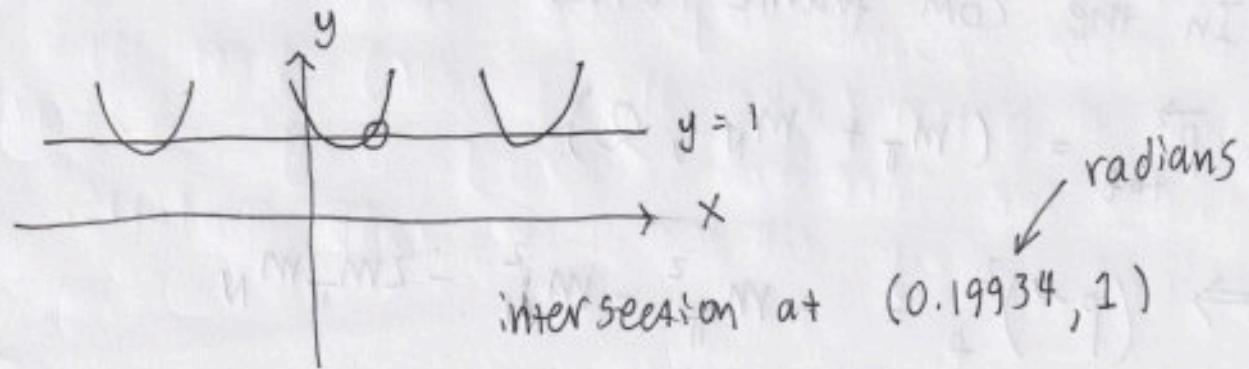
- Special Relativity is not violated in this case because this is just the apparent motion / information is not actually being transmitted at $>c$ values... 😊

C. For $V_{app} \approx 10c$, what is the largest possible value of θ ?

i.e. $10 = \frac{v \sin \theta}{1 - v \cos \theta}$ plot this as $\theta \rightarrow x$
 and $v \rightarrow y$ and ensure $y < 1$:

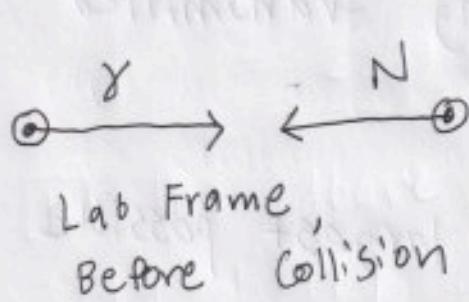
$$\Rightarrow 10 - 10y \cos x = y \sin x$$

$$\Rightarrow \frac{10}{\sin x + 10 \cos x} = y; \quad y < 1$$



$$\Rightarrow 0.199 \times \frac{180^\circ}{\pi} \approx 11.4^\circ \text{ is max possible } \theta \text{ for this configuration}$$

6 [a] Calculate the threshold energy of a nucleon N for it to undergo the reaction $\gamma + N \rightarrow N + \pi^0$ where γ is a CMB photon w/ energy kT where $T = 2.73$ kelvin. Assume the collision is head on & take $M_N = 938$ MeV and $M_{\pi} = 135$ MeV:



COM Frame,
After collision

(θ relative velocity s.t.
min energy required)

- \vec{p}^N is a conserved quantity (the momentum 4-vector) and $(p^N)^2$ {the square of 4-mom between frames}. Therefore, to solve relativistic kinematics problems like this we will compare $(p^N)^2_{\text{initial}} = (p^N)^2_{\text{final}}$

- In the COM frame after collision:

$$\vec{p}_{\text{tot}} = (m_{\pi} + m_N, \ell)$$

$$\Rightarrow (p^N)^2_p = -m_{\pi}^2 - m_N^2 - 2m_{\pi}m_N$$

In the lab frame before the collision,

$$\vec{P}_{\text{tot}} = \vec{P}_\gamma + \vec{P}_N$$

$$\Rightarrow (\vec{p}^\mu)_i^2 = P_\gamma^2 + P_N^2 + 2 \vec{P}_\gamma \cdot \vec{P}_N$$

Lorentz Invariants

In lab frame; $\vec{P}_\gamma = (\varepsilon, -\vec{\varepsilon})$

$\varepsilon = kT$

Since $E^2 = P^2 + m^2$ zero rest mass for photon

$$\Rightarrow |\vec{P}|_{\substack{\text{Spatial} \\ \text{photon}}} = \varepsilon$$

$$\Rightarrow P_\gamma^2 = -\varepsilon^2 + \varepsilon^2 = 0$$

Also in lab frame; $\vec{P}_N \approx (E_N, \vec{E}_N)$

Since we assume $|\vec{P}_{N, \text{spatial}}| \gg m_N$

$$\Rightarrow 2 \vec{P}_\gamma \cdot \vec{P}_N \approx 2(-\varepsilon E_N - \varepsilon E_N) = -4kT E_N$$

In rest frame of nucleon; $\vec{P}_N = (m_N, \vec{0})$

$$\Rightarrow P_N^2 = -m_N^2$$

• Putting this all together yields:

$$-\cancel{m_N^2} - m_{\pi}^2 - 2m_{\pi}m_N = \cancel{-m_N^2} - 4kT E_N$$
$$\Rightarrow E_N = \frac{m_{\pi}^2 + 2m_{\pi}m_N}{4kT}$$

$$\Rightarrow E_N \approx 2.89 \times 10^{20} \text{ eV}$$
$$= 2.89 \times 10^{11} \text{ GeV}$$

- b) The above result tells us that it is very likely any would collide with nucleons above this energy. It would then seem unlikely to find any freely traveling nucleons above this energy. This is called the Griesen-Zatsepin-Kuzmin GZK cutoff.

- c) Many examples of cosmic rays above the GZK cutoff have been found. These are puzzling, but if these cosmic rays are made of nucleons heavier than a proton (like He, Li, etc) then this would raise the threshold energy they could achieve before they collide with CMB photons.